



# Theoretical calculation of oxygen equilibrium isotope fractionation factors involving various $\text{NO}_y$ molecules, $\cdot\text{OH}$ , and $\text{H}_2\text{O}$ and its implications for isotope variations in atmospheric nitrate

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## Abstract

The oxygen stable isotope composition ( $\delta^{18}\text{O}$ ) of nitrogen oxides [ $\text{NO}_x$  = nitric oxide (NO) + nitrogen dioxide ( $\text{NO}_2$ )] and their oxidation products ( $\text{NO}_y$  =  $\text{NO}_x$  + nitric acid ( $\text{HNO}_3$ ) + particulate nitrate (p- $\text{NO}_3^-$ ) + nitrate radical ( $\text{NO}_3$ ) + dinitrogen pentoxide ( $\text{N}_2\text{O}_5$ ) + nitrous acid (HONO) + ...) have been shown to be a useful tool for inferring the proportion of  $\text{NO}_x$  that is oxidized by ozone ( $\text{O}_3$ ). However, isotopic fractionation processes may have an influence on  $\delta^{18}\text{O}$  of various  $\text{NO}_y$  molecules and other atmospheric O-bearing molecules pertinent to  $\text{NO}_x$  oxidation chemistry. Here we have evaluated the impacts of O isotopic exchange involving  $\text{NO}_y$  molecules, the hydroxyl radical ( $\cdot\text{OH}$ ), and water ( $\text{H}_2\text{O}$ ) using reduced partition function ratios ( ${}^{\circ}\beta$ ) calculated by hybrid density functional theory. Assuming atmospheric isotopic equilibrium is achieved between NO and  $\text{NO}_2$  during the daytime, and  $\text{NO}_2$ ,  $\text{NO}_3$ , and  $\text{N}_2\text{O}_5$  during the nighttime,  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  compositions were predicted for the major atmospheric nitrate formation pathways using our calculated exchange fractionation factors and isotopic mass-balance. Our equilibrium nitrate model predicts that various atmospheric nitrate formation pathways, including  $\text{NO}_2 + \cdot\text{OH} \rightarrow \text{HNO}_3$ ,  $\text{N}_2\text{O}_5 + \text{H}_2\text{O} + \text{surface} \rightarrow 2\text{HNO}_3$ , and  $\text{NO}_3 + \text{R} \rightarrow \text{HNO}_3 + \text{R}'$  will yield distinctive  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  compositions. Our calculated  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  compositions match well with previous atmospheric nitrate measurements, and will potentially help better understand the role oxidation chemistry plays on the N and O isotopic composition of atmospheric nitrate.

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## 1. INTRODUCTION

Stable isotopes are useful for understanding complex chemical systems such as the cycling and oxidation of

$\text{NO}_x$  [ $\text{NO}_x$  = nitric oxide (NO) + nitrogen dioxide ( $\text{NO}_2$ )]. For example, the oxygen (O) stable isotope composition of atmospheric nitrate, which includes nitric acid ( $\text{HNO}_3$ ), nitrate ( $\text{NO}_3^-$ ), and particulate nitrate (p- $\text{NO}_3^-$ ), indicates the proportion of  $\text{NO}_x$  that reacts with ozone ( $\text{O}_3$ ) during its oxidation (Michalski et al., 2003; Thiemens, 2006; Savarino et al., 2007; McCabe et al., 2007; Morin et al., 2008; Alexander et al., 2009). The O isotope composition

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of a sample is typically reported in delta notation ( $\delta$ ), which is expressed as the ratio of the heavy ( $^{17}\text{O}$  and  $^{18}\text{O}$ ) to the light ( $^{16}\text{O}$ ) isotope in a sample relative to the same ratio of an international standard (Vienna Standard Mean Ocean Water (VSMOW)):

$$\delta^x\text{O}(\text{‰}) = \left[ \left( \frac{[^x\text{O}/^{16}\text{O}]_{\text{sample}}}{(^x\text{O}/^{16}\text{O})_{\text{VSMOW}}} \right) - 1 \right] \times 1000 \quad (1)$$

where  $x$  represents the abundance of either  $^{17}\text{O}$  or  $^{18}\text{O}$ . Several studies have shown that atmospheric  $\text{O}_3$  has elevated  $\delta^{17}\text{O}$  and  $\delta^{18}\text{O}$  values relative to VSMOW and a strong mass-independent component (Thiemens and Heidenreich, 1983; Krankowsky et al., 1995; Johnston and Thiemens, 1997; Mauersberger et al., 2001; Vicars et al., 2012; Vicars and Savarino, 2014), which is quantified by  $\Delta^{17}\text{O}$  notation:

$$\Delta^{17}\text{O}(\text{‰}) = 1000 \ln \left[ 1 + \frac{\delta^{17}\text{O}}{1000} \right] - 0.52 \times 1000 \ln \left[ 1 + \frac{\delta^{18}\text{O}}{1000} \right] \quad (2)$$

These elevated  $\delta^{18}\text{O}$  and  $\Delta^{17}\text{O}$  signatures are transferred to atmospheric nitrate proportionally when  $\text{O}_3$  oxidizes  $\text{NO}_x$  (Thiemens, 2006; Savarino et al., 2008; Michalski and Bhattacharya, 2009; Berhanu et al., 2012). Thus, O isotopic analysis, particularly  $\Delta^{17}\text{O}$ , of atmospheric nitrate has been suggested as a useful proxy for assessing changes in  $\text{NO}_x$  oxidation and for evaluating long-term changes in the atmosphere's oxidation capacity (e.g. Michalski et al., 2003; Alexander et al., 2009; Alexander and Mickley, 2015).

There may be additional useful information in  $\delta^{18}\text{O}$  variations, because mass-dependent fractionation processes (MDFP) such as equilibrium or kinetic isotope effects will have a minimal impact on  $\Delta^{17}\text{O}$  but may induce significant  $\delta^{18}\text{O}$  variations. These fractionation processes may also influence the nitrogen (N) isotopic composition ( $\delta^{15}\text{N}(\text{‰}) = \left[ \frac{(^{15}\text{N}/^{14}\text{N})_{\text{sample}}}{(^{15}\text{N}/^{14}\text{N})_{\text{air}}} - 1 \right] \times 1000$ , where air  $\text{N}_2$  is the N isotopic reference) as  $\text{NO}_x$  is oxidized to atmospheric nitrate (e.g. Freyer et al., 1993; Vicars et al., 2013; Walters et al., 2016). Thus,  $\delta^{18}\text{O}$  and  $\delta^{15}\text{N}$  of atmospheric nitrate may be related, reflecting to some degree the fractionation processes responsible for the formation of atmospheric nitrate. However, few studies have quantitatively evaluated the impacts of  $\text{NO}_x$  oxidation fractionation processes on  $\delta^{18}\text{O}$  and their implications for possible  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  relationships of the atmospheric nitrate end product.

One such fractionation process that may play an important role on the  $\delta^{18}\text{O}$  and  $\delta^{15}\text{N}$  of atmospheric nitrate is isotopic equilibrium exchange or partial exchange. For example, it has been previously suggested that isotopic equilibrium between  $\text{NO}$  and  $\text{NO}_2$  has a significant impact on the  $\delta^{15}\text{N}$  of  $\text{HNO}_3$  produced during the daytime (Freyer et al., 1993; Riha, 2013; Vicars et al., 2013; Savarino et al., 2013; Walters et al., 2016). The same exchange processes may also generate useful  $\delta^{18}\text{O}$  signatures in  $\text{HNO}_3$ . Additionally, it has been suggested that  $\text{NO}_2$ , the nitrate radical ( $\text{NO}_3$ ), and dinitrogen pentoxide ( $\text{N}_2\text{O}_5$ ) may reach isotopic equilibrium during the nighttime (Amell and Daniels, 1952; Freyer, 1991; Walters et al., 2016). Isotopic equilibrium involving  $\text{NO}_2$ ,  $\text{NO}_3$ , and  $\text{N}_2\text{O}_5$  will affect partitioning of  $^{18}\text{O}$  and  $^{15}\text{N}$  between these molecules, which has

implications for variations in  $\delta^{18}\text{O}$  and  $\delta^{15}\text{N}$  of  $\text{HNO}_3$  produced at night. Thus, isotopic equilibrium exchange in addition to mass-balance considerations may have important implications for diurnal and seasonal variations in  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  compositions of atmospheric nitrate. In this case, knowing the isotopic equilibrium fractionation factors involving a variety of oxidized N molecules (denoted as  $\text{NO}_y = \text{NO}_x + \text{HNO}_3 + \text{p-NO}_3^- + \text{NO}_3 + \text{N}_2\text{O}_5 + \text{nitrous acid (HONO)} + \dots$ ) would be useful.

Prior studies have calculated O isotopic equilibrium exchange fractionation factors involving some  $\text{NO}_y$  molecules using measured vibrational spectroscopic data (Richet et al., 1977) or empirically determined interatomic force constants (Stern et al., 1968; Monse et al., 1969). However, O isotopic equilibrium fractionation factors have not been determined for many atmospherically relevant  $\text{NO}_y$  molecules such as  $\text{N}_2\text{O}_5$ ,  $\text{NO}_3$ , and halogen nitrates ( $\text{XONO}_2$ ). This is primarily due to the absence of spectroscopic data for the isotopologues of these molecules. Previously, we have used computational quantum chemistry methods to calculate N isotopic equilibrium fractionation factors involving some  $\text{NO}_y$  molecules (Walters and Michalski, 2015). Here we will extend that study and calculate O isotopic equilibrium exchange fractionation factors for singly substituted O isotopologues of numerous  $\text{NO}_y$  molecules. We will also calculate the fractionation factor associated with equilibrium isotopic exchange between the hydroxyl radical ( $\cdot\text{OH}$ ) and water ( $\text{H}_2\text{O}$ ) due to the importance this reaction has on the  $\delta^{18}\text{O}$  of atmospheric  $\cdot\text{OH}$  (Dubey et al., 1997; Lyons, 2001; Michalski et al., 2012). Knowing  $\delta^{18}\text{O}$  of  $\cdot\text{OH}$  is important because it will be transferred proportionally to the atmospheric nitrate end product from  $\cdot\text{OH}$  reaction with  $\text{NO}_2$  (Michalski et al., 2012). Our calculated O isotopic equilibrium fractionation factors will allow for the evaluation of the impact various exchange reactions have on  $\delta^{18}\text{O}$  values of these molecules and how it may be propagated into the atmospheric nitrate end product. Combining our O and N isotopic equilibrium fractionation factors will allow for the prediction of  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  arrays resulting from various atmospheric nitrate formation pathways.

## 2. METHODS AND THEORY

### 2.1. Calculation of equilibrium isotope exchange fractionation factors

The theoretical calculation of isotopic fractionation associated with isotope exchange equilibrium reactions has long been carried out in the harmonic approximation using partition functions (Urey, 1947; Bigeleisen and Mayer, 1947). In the Born–Oppenheimer, harmonic oscillator, and rigid-rotor approximations, the reduced partition function ratio (RPFR, commonly denoted as  ${}^x\beta$ ) for an oxygen isotopic pair is written as:

$$\text{RPFR} = {}^x\beta_{\text{har}} = \prod_i \frac{u_{2i}}{u_{1i}} \times \exp \left( \sum_i \frac{u_{1i} - u_{2i}}{2} \right) \times \prod_i \frac{1 - \exp(-u_{1i})}{1 - \exp(-u_{2i})} \quad (3)$$

where  $x$  and subscript 2 refer to one of the heavy isotopes ( $^{17}\text{O}$  or  $^{18}\text{O}$ ) and subscript 1 refers to  $^{16}\text{O}$ . The  $u_i = hc\omega_i/kT$  where  $h$  is Planck's constant,  $c$  is the speed of light,  $\omega_i$  is the harmonic frequency ( $\text{cm}^{-1}$ ) for each vibrational mode ( $i$ ),  $k$  is Boltzmann constant,  $T$  is temperature in Kelvin, and  $N$  refers to the number of normal mode vibrational frequencies. In Eq. (3), the Teller–Redlich rule was employed (Redlich, 1935), which converts translational and rotational motion into vibrational frequencies. Symmetry numbers are not considered, because symmetry in itself does not lead to isotopic enrichment during equilibrium (Bigeleisen and Mayer, 1947) but must be considered from a statistical perspective (Coulson, 1978; Pollak and Pechukas, 1978; Michalski and Bhattacharya, 2009).

The  ${}^x\beta$  values calculated using Eq. (3) are accurate in the harmonic oscillator and rigid-rotor approximations. However, corrections for anharmonicity, vibration–rotation coupling, centrifugal distortion, hindered internal rotation, and quantum mechanical correction to rotation have been shown to be important for calculating accurate  ${}^x\beta$  values for some molecules (Pennington and Kobe, 1954; Richet et al., 1977; Liu et al., 2010). Generally, for atoms other than H most of the corrections can be ignored as  ${}^x\beta$  is impacted by less than 0.01%. The main exception is anharmonic correction to the zero point energy (ZPE), which can play a significant role in the calculation of accurate  ${}^x\beta$  values (Liu et al., 2010). If accurate ZPEs are known or can be accounted for,  ${}^x\beta$  values can more accurately be calculated by the following (Liu et al., 2010):

$${}^x\beta_{\text{anhar}} = \frac{\exp[-\text{ZPE}_{\text{anhar},2}/kT]}{\exp[-\text{ZPE}_{\text{anhar},1}/kT]} \prod_i^N \left[ \frac{u_{2i}}{u_{1i}} \right] \times \left[ \frac{1 - \exp(-u_{1i})}{1 - \exp(-u_{2i})} \right] \quad (4)$$

where  $\text{ZPE}_{\text{anhar}}$  is the anharmonic corrected ground ZPE.

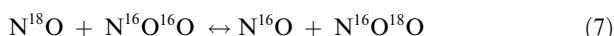
A general representation for an equilibrium isotope exchange reaction involving two different chemical compounds, A and B, is:



where subscripts 1 and 2 again refer to the light and heavy isotopologue, respectively. The equilibrium constant,  ${}^xK_{\text{A/B}}$ , which is also called the equilibrium isotope fractionation factor ( ${}^x\alpha_{\text{A/B}}$ ), can be obtained from the  ${}^x\beta$  values of A and B:

$${}^xK_{\text{A/B}} = {}^x\alpha_{\text{A/B}} = \frac{{}^x\beta_{\text{A}}}{{}^x\beta_{\text{B}}} \quad (6)$$

For example, using this notation, the equilibrium isotope exchange reaction between NO and  $\text{NO}_2$  involving the singly substituted  $^{18}\text{O}$  and  $^{16}\text{O}$  isotopologue pair is written as:



The equilibrium constant for this isotope exchange reaction is written as:

$${}^{18}K_{\text{NO}_2/\text{NO}} = {}^{18}\alpha_{\text{NO}_2/\text{NO}} = \frac{{}^{18}\beta_{\text{NO}_2}}{{}^{18}\beta_{\text{NO}}} \quad (8)$$

## 2.2. Computational chemistry methods

Using *ab initio* methods to obtain  $\beta$  and  $\alpha_{\text{A/B}}$  values for use in geochemical systems is a relatively recent advance (e.g., Driesner et al., 2000; Yamaji et al., 2001; Schauble et al., 2004, 2006; Anbar et al., 2005; Liu and Tossell, 2005; Otake et al., 2008; Liu et al., 2010). In this study, the Becke-3 parameter-Lee-Yang-Parr (B3LYP) hybrid density function theory (DFT) method (Lee et al., 1988; Becke, 1993) and Dunning correlation-consistent polarized valence triple  $\zeta$  (cc-pVTZ) basis set (Dunning, 1989) were used to calculate geometries (bond angles and bond lengths) and harmonic frequencies for the following molecules: NO,  $\text{NO}_2$ , HONO, HNO<sub>3</sub>,  $\text{N}_2\text{O}_5$ , nitrate ( $\text{NO}_3^-$ ), chlorine nitrate ( $\text{ClONO}_2$ ), nitrite ( $\text{NO}_2^-$ ), dinitrogen tetroxide ( $\text{N}_2\text{O}_4$ ), nitryl bromide ( $\text{BrNO}_2$ ), nitryl chloride ( $\text{ClNO}_2$ ), bromine nitrate ( $\text{BrONO}_2$ ),  $\cdot\text{OH}$ , and  $\text{H}_2\text{O}$ . We also calculated the geometry and vibrational frequencies of  $\text{NO}_3$  using the EDF2 DFT method (Lin et al., 2004) with the cc-pVTZ basis set (EDF2/cc-pVTZ) and the QCISD method with the cc-pVDZ basis set (QCISD/cc-pVDZ). These additional methods were used to compare the calculated vibrational frequencies of  $\text{NO}_3$  due to the difficulties this molecule presents for computational studies related to it being a polyatomic doublet radical with multireference character (e.g. Morris et al., 1990; Dutta et al., 2013). Calculations involving the B3LYP/cc-pVTZ and QCISD/cc-pVDZ methods were performed using the Gaussian09 program package revision D.01 (Frisch et al., 2009) on the Purdue Radon cluster (ITaP Research Computing, 2015). The EDF2/cc-pVTZ calculations were performed using the QChem 4.2 program package (Shao et al., 2015) on an IBM personal computer.

The masses of the most abundant isotopes of each element ( $^1\text{H}$ ,  $^{12}\text{C}$ ,  $^{14}\text{N}$ ,  $^{16}\text{O}$ ,  $^{35}\text{Cl}$ ,  $^{79}\text{Br}$ ) were used to calculate harmonic frequencies, and isotope effects were subsequently determined via single-atom substitutions of either  $^{17}\text{O}$  or  $^{18}\text{O}$  for each  $^{16}\text{O}$  position. Systematic model errors and anharmonic corrections to the ZPE were accounted for by applying a constant scale factor to the calculated harmonic ZPE, as recommended by Liu et al. (2010). We used a constant scale factor of 0.9787, recommended by Sinha et al. (2004) to correct calculated harmonic ZPE using B3LYP/cc-pVTZ to match experimental ZPEs. This scale factor was also used to correct the ZPE of  $\text{NO}_3$  calculated using EDF2/cc-pVTZ. The ZPE of  $\text{NO}_3$  calculated from QCISD/cc-pVDZ was scaled by 0.9776, based on a recommended value for QCISD/6-31G(d) (Scott and Radom, 1996).

## 3. RESULTS

### 3.1. Calculated ${}^x\beta$ values

Calculated geometries (Table S1), harmonic frequencies (Table S2), and ZPEs (Table S3) for the molecules of interest in this study are presented in the Supplementary data. The  $\text{NO}_3$  vibrational frequencies calculated using B3LYP/cc-pVTZ, EDF2/cc-pVTZ, QCISD/cc-pVDZ, and EOM-CCSD/aug-cc-pVDZ (Walters and Michalski, 2015)

Table 1

Comparison of calculated  $\text{NO}_3$   $\nu_i$  ( $\text{cm}^{-1}$ ) from different computational methods and experimental determined values. Values in parentheses represent the absolute difference between the calculated and experimental frequency for a particular vibrational mode.

	$\nu_1$ (asym bend)	$\nu_2$ (asym bend)	$\nu_3$ (umbrella)	$\nu_4$ (sym stretch)	$\nu_5$ (asym stretch)	$\nu_6$ (asym stretch)	Average error <sup>a</sup>
B3LYP/cc-pVTZ <sup>a</sup>	204.93 (160.7)	206.64 (159.0)	809.25 (46.2)	1107.09 (54.1)	1108.04 (8.0)	1129.61 (29.6)	76.3
EDF2/cc-pVTZ <sup>a</sup>	337.75 (27.9)	339.79 (25.8)	816.47 (53.4)	1142.11 (89.1)	1171.90 (71.9)	1174.35 (74.3)	57.1
QCISD/cc-pVDZ <sup>a</sup>	201.82 (163.8)	203.80 (161.8)	813.92 (50.8)	1144.00 (91.0)	1168.86 (68.9)	1169.08 (69.1)	100.9
EOM-CCSD/ aug-cc-pVDZ <sup>b</sup>	310.80 (54.8)	312.94 (52.7)	743.37 (19.7)	1103.92 (50.9)	1263.96 (164.0)	1264.34 (164.3)	84.4
Experiment	365.6 <sup>c</sup>	365.6 <sup>c</sup>	763.1 <sup>d</sup>	1053.0 <sup>e</sup>	1100.0 <sup>e</sup>	1100.0 <sup>e</sup>	

<sup>a</sup> Calculated from this work.

<sup>b</sup> Calculated previously from Walters and Michalski (2015).

<sup>c</sup> Beckers et al. (2009).

<sup>d</sup> Kim et al. (1992).

<sup>e</sup> Jacox and Thompson (2008).

<sup>\*</sup> Calculated as the average absolute deviation from experimental vibrational frequencies (i):  $\frac{1}{6} \sum_{i=1}^6 |(v_i)_{\text{calculated}} - (v_i)_{\text{experiment}}|$ .

were compared (Table 1) with experimental data (Kim et al., 1992; Jacox and Thompson, 2008; Beckers et al., 2009). Generally, it is found the most accurate  $\text{NO}_3$  vibrational frequencies are from the EDF2/cc-pVTZ method (Table 1). Using calculated frequencies,  $^{18}\beta$  values were calculated for each molecule over a temperature range from 150 to 450 K using Eq. (4). For the molecules that had non-equivalent O sites (i.e.  $\text{HNO}_3$ ,  $\text{N}_2\text{O}_5$ ,  $\text{ClONO}_2$ ,  $\text{HONO}$ , and  $\text{BrONO}_2$ ), two  $^{18}\beta$  values were calculated. These are designated as “c” to refer to the substitution of  $^{18}\text{O}$  in the central bridging O atom position (e.g.  $\text{HONO}^c = \text{H}^{18}\text{ONO}$ ) and “t” to refer to the substitution of  $^{18}\text{O}$  in the terminal O atom position (e.g.  $\text{HONO}^t = \text{HON}^{18}\text{O}$ ). For molecules with more than one terminal O atom, but with slight break in symmetry (e.g.  $\text{HNO}_3$ ,  $\text{BrONO}_2$ , and  $\text{ClONO}_2$ ), the geometric mean  $^{18}\beta$  at all terminal sites is reported. These terminal sites are nearly equivalent as their O  $^{18}\beta$  values were generally within 0.001.

Regression fit parameters calculated for  $^{18}\beta$  as a function of temperature, and  $^{18}\beta$  values at 298 K, are reported in Table 2. Calculated  $^{17}\beta$  regression fit parameters and values at 298 K can be found in the Supporting data (Table S4). At 298 K,  $^{18}\beta$  ranged from 1.0258 to 1.1110, increasing in the following order:  $\cdot\text{OH} < \text{H}_2\text{O} < \text{BrONO}_2^c < \text{ClONO}_2^c < \text{N}_2\text{O}_5^c < \text{HONO}^c < \text{NO}_3 < \text{HNO}_3^c < \text{NO}_3^- < \text{NO}_2 < \text{NO}_2^- < \text{N} < \text{N}_2\text{O}_4 < \text{BrNO}_2 < \text{ClNO}_2 < \text{HNO}_3^t < \text{HONO}^t < \text{BrONO}_2^t < \text{ClONO}_2^t < \text{N}_2\text{O}_5^t$  (Table 2). The magnitude of  $^{18}\beta$  appears to be related to the number of bonds and the electronegativity of the oxygen bond-partner and increased in the following order: 1 hydrogen < 2 hydrogen < 1 halogen + 1 nitrogen < 2 nitrogen < 1 hydrogen + 1 nitrogen < 1 nitrogen (Table 2). This ordering of atoms that O is directly bonded to may provide a rough *a priori* way to arrange  $^{18}\beta$  values based solely on molecular structure.

Fig. 1 compares our calculated  $^{18}\beta$  values for  $\text{NO}$ ,  $\text{NO}_2$ ,  $\text{H}_2\text{O}$ , and  $\text{HNO}_3^c$  with those previously reported (Monse et al., 1969; Richet et al., 1977; Liu et al., 2010) as well as calculated using experimental data (Michalski, 2003; Michalski et al., 2004). Overall, our calculated  $^{18}\beta$  values agree with those previously reported (Fig. 1). For example, at 300 K, we calculated  $^{18}\beta_{(\text{NO})}$  to be 1.0995, which is close

to previously calculated values of 1.0973 (Richet et al., 1977) and 1.0963 (Monse et al., 1969). Additionally, our calculated  $^{18}\beta_{(\text{NO}_2)}$  at 300 K is 1.0909, which is close to the value of 1.0899 previously calculated by Monse et al. (1969) and the value of 1.0867 calculated from empirically determined interatomic force constants (Michalski, 2003) and experimental ZPEs (Michalski et al., 2004). For  $\text{H}_2\text{O}$ , we calculated  $^{18}\beta_{(\text{H}_2\text{O})}$  at 300 K to be 1.0635, which is close to values previously calculated of 1.0632 (Liu et al., 2010) and 1.0630 (Richet et al., 1977). Finally, for  $\text{HNO}_3^c$ , we calculated  $^{18}\beta_{(\text{HNO}_3^c)}$  to be 1.0862 at 300 K, which is close to the value of 1.0860 reported by Monse et al. (1969). Fig. 2 compares the  $^{18}\beta$  calculated for  $\text{NO}_3$  using B3LYP/cc-pVTZ, EDF2/cc-pVTZ, and QCISD/cc-pVDZ methods and indicates general agreement. For example, at 300 K,  $^{18}\beta$  for  $\text{NO}_3$  is calculated to be 1.0775, 1.08135, and 1.0796 for B3LYP/cc-pVTZ, EDF2/cc-pVTZ, and QCISD/cc-pVDZ, respectively. However, since EDF2/cc-pVTZ calculated the most accurate  $\text{NO}_3$  vibrational frequencies of the considered computational methods (Table 1), we used the EDF2/cc-pVTZ calculated  $^{18}\beta$  values for  $^{18}\alpha_{A/B}$  calculations involving  $\text{NO}_3$ .

### 3.2. Calculated $^{18}\alpha_{A/B}$ values

Using our calculated  $^{18}\beta$  values,  $^{18}\alpha_{A/B}$  were calculated for the following exchange reactions (A/B):  $\text{NO}/\text{NO}_2$ ,  $\text{NO}_2/\text{NO}_3$ ,  $\text{N}_2\text{O}_5^{\text{geo}}/\text{NO}_2$ ,  $\text{N}_2\text{O}_4/\text{NO}_2$ ,  $\text{BrNO}_2/\text{NO}_2$ ,  $\text{ClNO}_2/\text{NO}_2$ ,  $\text{BrONO}_2^t/\text{NO}_2$ ,  $\text{ClONO}_2^t/\text{NO}_2$ ,  $\text{NO}_3^-/\text{H}_2\text{O}$ , and  $\text{H}_2\text{O}/\text{OH}$  over the temperature range of 150–450 K, where “geo” refers to the geometric mean  $^{18}\beta$ . These exchange reactions were chosen due to their common molecular structure (e.g.  $\text{NO}$ ,  $\text{NO}_2$ , or  $\text{OH}$  subunits), and therefore, possibility that O isotopic exchange might occur at a rate comparable to the atmospheric lifetime of these compounds ( $\sim 1$  week). Few studies have experimentally determined  $\alpha_{A/B}$  involving the considered molecules, limiting our ability to evaluate our calculated  $\alpha_{A/B}$  values. One study that we are aware of involves the O isotopic exchange between  $\text{NO}_3^-$  and  $\text{H}_2\text{O}$  (Böhlke et al., 2003), in which  $\alpha_{\text{NO}_3^-/\text{H}_2\text{O}}$  was determined to be 1.0215 and 1.0131 at 295 K and

Table 2

Calculated  $^{18}\beta$  regression coefficients as a function of temperature ( $150\text{ K} \leq T \leq 450\text{ K}$ ) and  $^{18}\beta$  (298 K) values for  $\text{NO}_y$  molecules,  $\cdot\text{OH}$ , and  $\text{H}_2\text{O}$ .

$$1000(^{18}\beta - 1) = \frac{A}{T^4} \times 10^{10} + \frac{B}{T^3} \times 10^8 + \frac{C}{T^2} \times 10^6 + \frac{D}{T} \times 10^4$$

	A	B	C	D	$^{18}\beta$ (298 K)
$\cdot\text{OH}$	1.1036	-1.9577	1.2889	0.51636	1.0258
$\text{H}_2\text{O}$	3.2070	-5.7294	3.9525	1.1030	1.0639
$\text{BrONO}_2^c$	2.1489	-6.1744	8.3297	-0.22623	1.0656
$\text{ClONO}_2^c$	2.6464	-7.0231	8.6922	-0.22784	1.0671
$\text{N}_2\text{O}_5^c$	2.6369	-6.9058	8.6047	-0.20778	1.0672
$\text{HONO}^c$	3.3651	-7.2949	7.0520	0.51606	1.0734
$\text{NO}_3^*$	6.7853	-13.499	11.279	-0.08248	1.0818
$\text{HNO}_3^c$	3.9484	-8.8137	9.0903	0.38699	1.0871
$\text{NO}_3^-$	6.3617	-13.388	12.124	-0.10652	1.0904
$\text{NO}_2$	7.8475	-15.028	11.690	0.20838	1.0918
$\text{NO}_2^-$	7.8614	-15.049	11.700	0.20718	1.0918
$\text{NO}$	8.1294	-14.507	10.124	0.93160	1.1008
$\text{N}_2\text{O}_4$	7.7029	-14.916	12.695	0.21039	1.1034
$\text{BrNO}_2$	7.9945	-15.453	13.127	0.19084	1.1060
$\text{ClNO}_2$	8.0617	-15.674	13.396	0.19632	1.1084
$\text{HNO}_3^t$	7.9932	-15.922	13.776	0.10526	1.1086
$\text{HONO}^t$	8.5480	-16.103	12.686	0.50745	1.1099
$\text{BrONO}_2^t$	8.0302	-15.747	13.703	0.14583	1.1099
$\text{ClONO}_2^t$	8.0808	-15.824	13.748	0.16979	1.1110
$\text{N}_2\text{O}_5^t$	8.1801	-15.985	13.792	0.17181	1.1110

“c” and “t” correspond to the substitution of  $^{18}\text{O}$  at the bridging or terminal O atom position, respectively for a particular molecule (e.g.  $\text{HONO}^c = \text{H}^{18}\text{ONO}$ ,  $\text{HONO}^t = \text{HON}^{18}\text{O}$ ).

Typical misfit in the regression is  $<0.01\%$ .

\* All  $^{18}\beta$  values calculated using B3LYP/cc-pVTZ, except for  $\text{NO}_3$ , which was calculated using EDF2/cc-pVTZ.

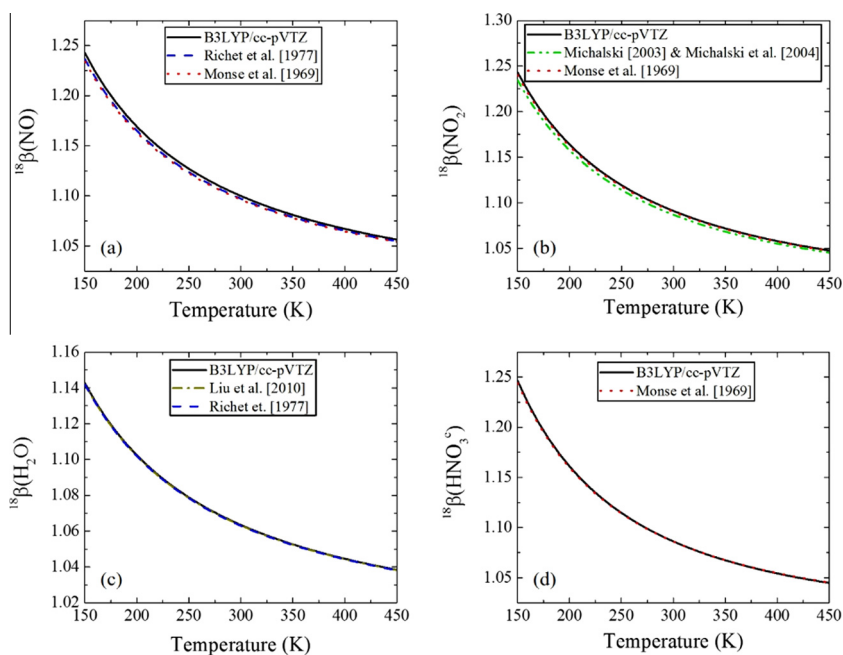


Fig. 1. Comparison of calculated  $^{18}\beta$  values using B3LYP/cc-pVTZ with those computed in previous studies for (a)  $\text{NO}$  (Monse et al., 1969; Richet et al., 1977), (b)  $\text{NO}_2$  (Monse et al., 1969; Michalski, 2003; Michalski et al., 2004), (c)  $\text{H}_2\text{O}$  (Richet et al., 1977; Liu et al., 2010), and (d)  $\text{HNO}_3$  (Monse et al., 1969).

373 K, respectively. These experimentally determined values agree reasonably well with our calculated values of 1.0256 and 1.0126 at 295 K and 373 K, respectively.

Table 3 presents the calculated  $^{18}\alpha_{A/B}$  regression fit coefficients for these exchange reactions sorted in order of increasing  $^{18}\alpha_{A/B}$  at 298 K. Overall, the calculated  $^{18}\alpha_{A/B}$

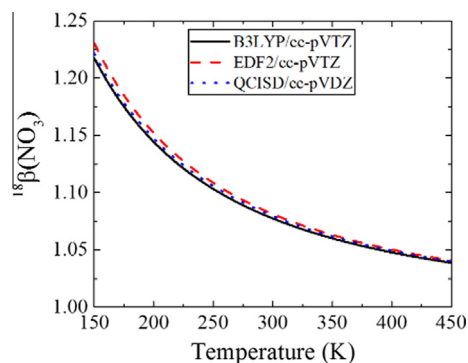


Fig. 2. Comparison of calculated  $^{18}\beta$  for  $\text{NO}_3$  using various computational methods that include B3LYP/cc-pVTZ, EDF2/cc-pVTZ, and QCISD/cc-pVDZ.

values ranged from 1.0082 to 1.0371 at 298 K (Table 3) and increased in the following order:  $\text{NO}/\text{NO}_2 < \text{NO}_2/\text{NO}_3 < \text{N}_2\text{O}_5^{\text{geo}}/\text{NO}_2 < \text{N}_2\text{O}_4/\text{NO}_2 < \text{BrNO}_2/\text{NO}_2 < \text{ClNO}_2/\text{NO}_2 < \text{BrONO}_2/\text{NO}_2 < \text{ClONO}_2/\text{NO}_2 < \text{NO}_3^-/\text{H}_2\text{O} < \text{H}_2\text{O}/\text{OH}$  (Table 3). This trend indicates that at the most exchangeable sites (i.e. terminal O atoms),  $^{18}\text{O}$  will generally partition into  $\text{NO}_y$  molecules other than  $\text{NO}$  and  $\text{NO}_2$ . This is true for all considered equilibrium exchange reactions except for  $\text{NO}_2/\text{NO}_3$ , in which  $^{18}\text{O}$  is found to prefer  $\text{NO}_2$ . Additionally, our calculations indicate that the exchange between  $\text{H}_2\text{O}/\text{OH}$  will result in atmospheric  $^{\text{OH}}$  depleted in  $^{18}\text{O}$  relative to local  $\text{H}_2\text{O}$  vapor by approximately  $-35\%$  at 298 K.

Fig. 3 displays the isotopic equilibrium exchange enrichment factors ( $^{18}\varepsilon_{\text{A/B}} = 1000(^{18}\alpha_{\text{A/B}} - 1)$ ) for the following five isotopic equilibrium exchange reactions:  $\text{NO}/\text{NO}_2$ ,  $\text{NO}_2/\text{NO}_3$ ,  $\text{N}_2\text{O}_5^{\text{geo}}/\text{NO}_2$ ,  $\text{ClNO}_2/\text{NO}_2$ , and  $\text{ClONO}_2/\text{NO}_2$ . Generally, as temperature increases, it is observed that  $^{18}\varepsilon_{\text{A/B}}$  decreases towards zero as expected for equilibrium isotope effects in the high temperature limit. This was the observed trend for all of the considered isotopic exchange reactions except for  $\text{NO}/\text{NO}_2$ , in which  $\alpha_{\text{A/B}}$  switches from less than 1 to greater than 1 at approximately 152.5 K. Fig. 3 also shows  $\Delta^{17}\text{O}$  arising from the considered isotopic

equilibrium exchange reactions where  $\Delta^{17}\text{O}$  was calculated according to the following:

$$\Delta^{17}\text{O}(\text{‰}) = 1000 \ln [1 + (^{17}\alpha_{\text{A/B}} - 1)] - 0.52 \times 1000 \ln [1 + (^{18}\alpha_{\text{A/B}} - 1)] \quad (9)$$

For the considered exchange reactions,  $\Delta^{17}\text{O}$  was found to minimally deviate from 0‰ (Fig. 3), ranging from 0.02 to 0.18‰ (Fig. 3) at 298 K.

#### 4. DISCUSSION

Overall, our  $^{18}\alpha_{\text{A/B}}$  calculations show that if isotopic equilibrium is achieved, it will influence the partitioning of  $^{18}\text{O}$  between the considered molecules, without inducing a significant  $\Delta^{17}\text{O}$  signature, as expected for a MDFP (Fig. 3). Therefore, while  $\Delta^{17}\text{O}$  of  $\text{NO}_y$  molecules should approximately represent the O mass-balance of precursor molecules (i.e.  $\text{O}_3$ ,  $\text{H}_2\text{O}$ ,  $\text{O}_2$ ),  $\delta^{18}\text{O}$  may reflect both O mass-balance and possible influences from MDFP such as isotopic equilibrium exchange. Thus,  $\delta^{18}\text{O}$  of  $\text{NO}_y$  molecules may provide additional information about  $\text{NO}_x$  oxidation pathways that is not observable in  $\Delta^{17}\text{O}$  alone. Taking into account N and O isotopic mass-balance and assuming daytime isotopic equilibrium between  $\text{NO}$  and  $\text{NO}_2$  (Freyer et al., 1993; Riha, 2013; Vicars et al., 2013; Savarino et al., 2013; Walters et al., 2016) and nighttime isotopic equilibrium between  $\text{NO}_2$ ,  $\text{NO}_3$ , and  $\text{N}_2\text{O}_5$  (Amell and Daniels, 1952; Freyer, 1991; Walters et al., 2016),  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  compositions may be estimated for various  $\text{HNO}_3$  formation pathways.

##### 4.1. Predicted daytime $\text{HNO}_3$ $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$ compositions

During the daytime,  $\text{NO}_x$  exists in a photochemical steady state in which  $\text{NO}$  is oxidized to  $\text{NO}_2$ , which may photolyze back to  $\text{NO}$  leading to the production of  $\text{O}_3$  (Leighton, 1961).



Table 3

Calculated  $^{18}\alpha_{\text{A/B}}$  regression coefficients as a function of temperature ( $150 \text{ K} \leq T \leq 450 \text{ K}$ ) and  $^{18}\alpha_{\text{A/B}}$  (298 K) values for O isotopic equilibrium exchange reactions.

	$1000(^{18}\alpha_{\text{A/B}} - 1) = \frac{A}{T^2} \times 10^{10} + \frac{B}{T^3} \times 10^8 + \frac{C}{T^4} \times 10^6 + \frac{D}{T} \times 10^4$				
	A	B	C	D	$^{18}\alpha_{\text{A/B}}(298 \text{ K})$
$\text{NO}/\text{NO}_2$	−0.04129	1.1605	−1.8829	0.74723	1.0082
$\text{NO}_2/\text{NO}_3$	1.03163	−1.38703	0.24875	0.3082	1.0092
$\text{N}_2\text{O}_5^{\text{geo}}/\text{NO}_2$	−0.54136	0.13073	1.2477	−0.1272	1.0096
$\text{N}_2\text{O}_4/\text{NO}_2$	−0.02333	−0.36888	1.0789	−0.00256	1.0106
$\text{BrNO}_2/\text{NO}_2$	0.29493	−0.96875	1.5039	−0.02007	1.0130
$\text{ClNO}_2/\text{NO}_2$	0.41012	−1.2744	1.7759	−0.01414	1.0152
$\text{BrONO}_2/\text{NO}_2$	0.43126	−1.4765	2.1167	−0.06681	1.0166
$\text{ClONO}_2/\text{NO}_2$	0.48353	−1.5568	2.1511	−0.04177	1.0176
$\text{NO}_3^-/\text{H}_2\text{O}$	3.6280	−8.6540	8.2763	−1.1983	1.0249
$\text{H}_2\text{O}/\text{OH}$	2.1137	−3.8026	2.5653	0.59410	1.0371

“t” corresponds to the substitution of  $^{18}\text{O}$  at the terminal O atom position for a particular molecule (e.g.  $\text{BrONO}_2^{\text{t}} = \text{BrONO}^*\text{O}$ ) and “geo” refers to the geometric mean  $^{18}\beta$  value for central and terminal mono-substituted  $^{18}\text{O}$  of  $\text{N}_2\text{O}_5$

Typical misfit in the regression is  $<0.01\%$ .

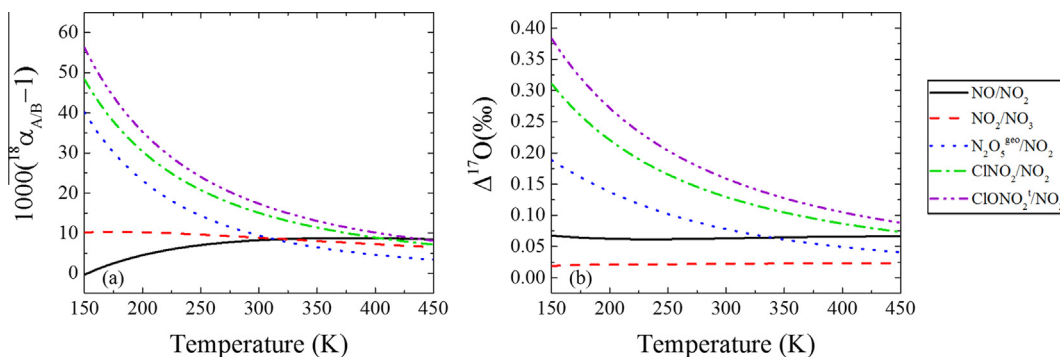


Fig. 3. (a) Calculated O isotopic exchange enrichment factors ( $1000(\alpha_{A/B}^{18} - 1)$ ) and (b) calculated O isotopic exchange induced mass-independence ( $\Delta^{17}\text{O}(\text{‰}) = 1000\ln(\alpha_{A/B}^{17}) - 0.52 * 1000\ln(\alpha_{A/B}^{18}) - 1$ ) for the following exchange reactions: NO/NO<sub>2</sub>, NO<sub>2</sub>/NO<sub>3</sub>, N<sub>2</sub>O<sub>5</sub><sup>860</sup>/NO<sub>2</sub>, ClNO<sub>2</sub>/NO<sub>2</sub>, and ClONO<sub>2</sub>/NO<sub>2</sub>.

The photochemical cycling between NO–NO<sub>2</sub>–O<sub>3</sub> is rapid, and prior experimental investigations of this cycling has suggested that O isotopic equilibrium is achieved between O<sub>3</sub> and NO<sub>x</sub> (Michalski et al., 2014). Based on a photochemical NO<sub>x</sub>–O<sub>3</sub> isotope equilibrium model,  $\delta^{18}\text{O}$ –NO<sub>x</sub> has been estimated to have a value of  $117 \pm 5\text{‰}$  relative to VSMOW (Michalski et al., 2014). While  $\delta^{18}\text{O}$ –NO<sub>x</sub> is assumed to reflect its photochemical equilibration with O<sub>3</sub>,  $\delta^{15}\text{N}$ –NO<sub>x</sub> should be related to local NO<sub>x</sub> source emissions. Recently, a mass-balance  $\delta^{15}\text{N}$ –NO<sub>x</sub> model across the contiguous United States has been developed that suggests  $\delta^{15}\text{N}$ –NO<sub>x</sub> should typically range from  $-15$ – $0\text{‰}$  (Walters et al., 2015). Exceptions include regions dominated by soil emissions (e.g. Great Plains during the summer) or coal-fired power plant emissions, in which  $\delta^{15}\text{N}$ –NO<sub>x</sub> is estimated to range from  $-31$ – $-21\text{‰}$  and  $5$ – $9\text{‰}$ , respectively (Walters et al., 2015).

While the isotopic composition of NO<sub>x</sub> is approximately known, the isotopic composition of its components, NO and NO<sub>2</sub>, may be altered relative to NO<sub>x</sub> due to isotopic exchange. The rate of isotopic exchange between NO and NO<sub>2</sub> is rapid ( $k = 8.14 \times 10^{-14} \text{ cm}^3 \text{ molecules}^{-1} \text{ s}^{-1}$  at 298 K; Sharma et al., 1970) and may lead to the partitioning of <sup>18</sup>O between NO and NO<sub>2</sub> similar to exchange previously observed for <sup>15</sup>N (Freyer et al., 1993; Walters et al., 2016). The resulting  $\delta^{18}\text{O}$  and  $\delta^{15}\text{N}$  of the NO and NO<sub>2</sub> will depend on the fraction of NO and NO<sub>2</sub> relative to the total NO<sub>x</sub> (i.e. [NO]/[NO<sub>x</sub>] and [NO<sub>2</sub>]/[NO<sub>x</sub>]) and the temperature dependent equilibrium exchange fractionation factors (Freyer et al., 1993; Walters et al., 2016). The impact of this exchange on  $\delta^{15}\text{N}$  of NO and NO<sub>2</sub> has been previously derived (Freyer et al., 1993; Walters et al., 2016), which we have adapted to also include  $\delta^{18}\text{O}$ , as the following:

$$\delta^x - \text{NO}_2(\text{‰}) = 1000 \left[ \frac{(\alpha_{\text{NO}_2/\text{NO}}^x - 1)(1 - f_{\text{NO}_2})}{(1 - f_{\text{NO}_2}) + (\alpha_{\text{NO}_2/\text{NO}}^x \cdot f_{\text{NO}_2})} \right] + \delta^x - \text{NO}_x(\text{‰}) \quad (10)$$

$$\delta^x - \text{NO}(\text{‰}) = \frac{(\delta^x - \text{NO}_x(\text{‰})) - (f_{\text{NO}_2} \cdot \delta^x - \text{NO}_2(\text{‰}))}{(1 - f_{\text{NO}_2})} \quad (11)$$

where  $x$  represents either <sup>18</sup>O or <sup>15</sup>N and  $f_{\text{NO}_2}$  is the fraction of NO<sub>2</sub> to the total NO<sub>x</sub>. During the daytime, near equal concentrations of NO and NO<sub>2</sub> can occur (e.g. Vicars et al., 2013; Walters et al., 2016) due to the emission of NO (e.g. Gao, 2007) and NO<sub>2</sub> photolysis (Leighton, 1961). Therefore, during the daytime,  $\delta^{18}\text{O}$  and  $\delta^{15}\text{N}$  of NO and NO<sub>2</sub> will likely reflect a complex function of the  $\delta^{18}\text{O}$  and  $\delta^{15}\text{N}$  of NO<sub>x</sub> and the NO<sub>x</sub> equilibrium isotopic fractionation factors. If the isotopic composition of daytime NO<sub>2</sub> is altered relative to NO<sub>x</sub>, it would have important consequences for  $\delta^{18}\text{O}$  and  $\delta^{15}\text{N}$  of atmospheric nitrate produced during daylight hours, because it is primarily formed through the reaction between NO<sub>2</sub> and photochemically produced  $\cdot\text{OH}$ :



We will refer to HNO<sub>3</sub> produced through R4 as HNO<sub>3</sub>(1) in the following discussion. Assuming no kinetic isotope fractionation associated with R4 (this will be discussed in a forth coming paper), the  $\delta^x$  of HNO<sub>3</sub>(1) can be calculated using mass-balance as the following:

$$\delta^{15}\text{N} - \text{HNO}_3(1)(\text{‰}) = \delta^{15}\text{N} - \text{NO}_2(\text{‰}) \quad (12)$$

$$\delta^{18}\text{O} - \text{HNO}_3(1)(\text{‰}) = \frac{2}{3}(\delta^{18}\text{O} - \text{NO}_2(\text{‰})) + \frac{1}{3}(\delta^{18}\text{O} - \cdot\text{OH}(\text{‰})) \quad (13)$$

This model indicates that the N and O isotopic composition of HNO<sub>3</sub>(1) can be approximately calculated if  $\delta^x$ –NO<sub>2</sub> and  $\delta^{18}\text{O}$ – $\cdot\text{OH}$  are known. Using this proposed model, we calculated daytime  $\delta^{15}\text{N}$ –NO<sub>2</sub> according to Eq. (10) assuming local  $\delta^{15}\text{N}$ –NO<sub>x</sub> to range from  $-15$ – $0\text{‰}$  (Walters et al., 2015) and using N isotopic exchange fractionation factors for NO<sub>2</sub>/NO from our previous study (Walters and Michalski, 2015). These NO<sub>2</sub>/NO exchange fractionation factors were adapted to include corrections for ZPE anharmonicity (adapted  $^{15}\beta$  and  $^{15}\alpha_{\text{NO}_2/\text{NO}}$  values can be found in Tables S5 and S6). Predicted daytime  $\delta^{18}\text{O}$ –NO<sub>2</sub> values, relative to VSMOW, were calculated from Eq. (10) assuming  $\delta^{18}\text{O}$ –NO<sub>x</sub> to range from  $112$ – $122\text{‰}$  (Michalski et al., 2014) and using O isotopic NO<sub>2</sub>/NO exchange fractionation factors calculated in this

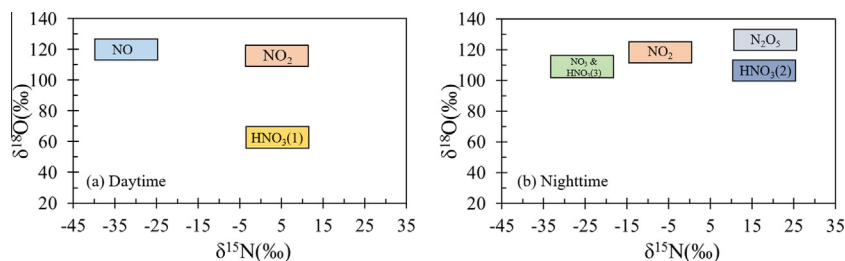


Fig. 4. Estimated  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  compositions of several  $\text{NO}_y$  molecules assuming (a) daytime isotopic equilibrium between  $\text{NO}/\text{NO}_2$  with  $f_{\text{NO}_2} = 0.7$  and (b) nighttime isotopic equilibrium between  $\text{NO}_2$ ,  $\text{NO}_3$ , and  $\text{N}_2\text{O}_5$ . Where  $\text{HNO}_3(1)$ ,  $\text{HNO}_3(2)$ , and  $\text{HNO}_3(3)$  represent various  $\text{HNO}_3$  production pathways that include  $\text{NO}_2 + \cdot\text{OH} \rightarrow \text{HNO}_3(1)$ ,  $\text{N}_2\text{O}_5 + \text{H}_2\text{O} + \text{surface} \rightarrow 2\text{HNO}_3(2)$ , and  $\text{NO}_3 + \text{R} \rightarrow \text{HNO}_3(3) + \text{R}$ .

study. The  $\delta^{18}\text{O}$  of  $\cdot\text{OH}$  should to a first order approximation reflect the  $\delta^{18}\text{O}$  of local water vapor and the temperature dependent fractionation factor associated with equilibrium  $\cdot\text{OH}/\text{H}_2\text{O}$  exchange, since the atmospheric concentration of  $\cdot\text{OH}$  is orders of magnitude lower than  $\text{H}_2\text{O}$  (Michalski et al., 2012):

$$\delta^{18}\text{O} - \cdot\text{OH}(\text{‰}) = \delta^{18}\text{O} - \text{H}_2\text{O}(\text{‰}) + 1000(\alpha_{\text{OH}/\text{H}_2\text{O}} - 1) \quad (14)$$

Therefore,  $\delta^{18}\text{O}$  of  $\cdot\text{OH}$  can be approximated using an estimated tropospheric water vapor  $\delta^{18}\text{O}$  range of  $-25$ – $0\text{‰}$  (typical for mid-latitudes Michalski et al., 2012) and O isotopic  $\cdot\text{OH}/\text{H}_2\text{O}$  exchange fractionation factors calculated in this study.

The predicted  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  ranges of  $\text{NO}$ ,  $\text{NO}_2$ , and  $\text{HNO}_3(1)$  at 300 K with  $f_{\text{NO}_2} = 0.70$  exhibit significant variation (Fig. 4a). The  $\delta^{15}\text{N}$  of  $\text{HNO}_3(1)$  and  $\text{NO}_2$  are calculated to range from  $-4.2$ – $10.8\text{‰}$ , while  $\delta^{15}\text{N}$  of  $\text{NO}$  is calculated to range from  $-40.2$ – $25.2\text{‰}$ . The positive shift in  $\delta^{15}\text{N}$  of  $\text{HNO}_3(1)$  and  $\text{NO}_2$  and the negative shift in  $\delta^{15}\text{N}$  of  $\text{NO}$  relative to the assumed  $\delta^{15}\text{N}$ – $\text{NO}_x$  value ( $-15$ – $0\text{‰}$ ) is a direct result of the N isotopic exchange between  $\text{NO}$  and  $\text{NO}_2$  that favors the partitioning of  $^{15}\text{N}$  into  $\text{NO}_2$  ( $^{15}\alpha_{\text{NO}_2/\text{NO}} = 1.0370$  at 300 K). The  $\delta^{18}\text{O}$  of  $\text{NO}$ ,  $\text{NO}_2$ , and  $\text{HNO}_3(1)$  are estimated to range from  $115.3$ – $125.3\text{‰}$ ,  $109.5$ – $119.5\text{‰}$ , and  $52.8$ – $67.8\text{‰}$  at 300 K (Fig. 4a). The positive shift in  $\delta^{18}\text{O}$  of  $\text{NO}$  and negative shift in  $\text{NO}_2$  relative to the assumed initial  $\delta^{18}\text{O}$ – $\text{NO}_x$  ( $112$ – $122\text{‰}$ ) is a result of their O isotopic exchange that favors the formation of  $^{18}\text{O}$  in  $\text{NO}$  ( $^{18}\alpha_{\text{NO}_2/\text{NO}} = 0.992$  at 300 K). The estimated  $\delta^{18}\text{O}$ – $\text{HNO}_3(1)$  shift of  $-56.7\text{‰}$  relative to  $\text{NO}_2$  is a result of the  $1/3$  O contribution from  $\cdot\text{OH}$  that is predicted to have a  $\delta^{18}\text{O}$  range of  $-60.5$ – $-35\text{‰}$  due to its equilibration with  $\text{H}_2\text{O}$  at 300 K ( $^{18}\alpha_{\text{OH}/\text{H}_2\text{O}} = 0.9645$  at 300 K).

The impact of varying  $f_{\text{NO}_2}$  values on  $\text{NO}$  and  $\text{NO}_2$  isotopic exchange and its influence on  $\delta^{15}\text{N}$  and  $\delta^{18}\text{O}$  of  $\text{NO}$ ,

$\text{NO}_2$ , and  $\text{HNO}_3(1)$  are displayed in Table 4. Generally, it is observed that  $\delta^{15}\text{N}$  is highly sensitive to  $f_{\text{NO}_2}$ , but  $\delta^{18}\text{O}$  is not. For example, from  $f_{\text{NO}_2} = 1$  to  $f_{\text{NO}_2} = 0.55$ , the shift in  $\delta^{15}\text{N}$  of  $\text{NO}_2$  and  $\text{HNO}_3(1)$  is  $16.3\text{‰}$ , while the change in  $\delta^{18}\text{O}$  of  $\text{NO}_2$  and  $\text{HNO}_3(2)$  is  $-3.8\text{‰}$  and  $-2.5\text{‰}$ , respectively (Table 4). The degree of sensitivity of  $\delta^{15}\text{N}$  and  $\delta^{18}\text{O}$  in  $\text{HNO}_3(1)$  to  $f_{\text{NO}_2}$  is a result of the magnitude of the  $\text{NO}_2/\text{NO}$  equilibrium fractionation factors, in which N fractionation is fairly large ( $^{15}\alpha_{\text{NO}_2/\text{NO}} = 1.0370$  at 300 K), while O fractionation is relatively minor ( $^{18}\alpha_{\text{NO}_2/\text{NO}} = 0.992$  at 300 K). This isotope equilibrium mechanism predicts that  $\delta^{15}\text{N}$ – $\text{HNO}_3(1)$  will be sensitive to diurnal fluctuations in  $f_{\text{NO}_2}$ , which is a phenomenon that has been previously observed (Vicars et al., 2013). Additionally, this mechanism indicates that since O fractionation between  $\text{NO}$  and  $\text{NO}_2$  is relatively minor,  $\delta^{18}\text{O}$ – $\text{HNO}_3(1)$  is approximately equal to the O isotope mass-balance between  $\text{NO}_x$ – $\text{O}_3$  (in photochemical equilibrium) and  $\cdot\text{OH}$ .

In addition to  $f_{\text{NO}_2}$  dependence, the isotopic composition of  $\text{HNO}_3(1)$  will also have a temperature dependence that is related to the temperature dependent  $\alpha_{\text{NO}_2/\text{NO}}$  and  $\alpha_{\text{OH}/\text{H}_2\text{O}}$  values. As temperature decreases,  $^{15}\alpha_{\text{NO}_2/\text{NO}}$  increases (Walters and Michalski, 2015; Walters et al., 2016), and this will result in  $\text{HNO}_3(1)$  having a higher  $\delta^{15}\text{N}$  value relative to the  $\delta^{15}\text{N}$ – $\text{NO}_x$  for a fixed  $f_{\text{NO}_2}$  value. Additionally, while  $^{18}\alpha_{\text{NO}_2/\text{NO}}$  is relatively insensitive to temperature (Fig. 2),  $^{18}\alpha_{\text{OH}/\text{H}_2\text{O}}$  values change significantly with temperature (i.e.  $^{18}\alpha_{\text{OH}/\text{H}_2\text{O}} = 0.9494$  and  $0.9682$  at 220 K and 330 K, respectively). Therefore, the  $\delta^{18}\text{O}$ – $\text{HNO}_3(1)$  formed at lower temperatures will have a lower  $\delta^{18}\text{O}$  that is primarily the result of the temperature dependence of isotopic exchange between  $\cdot\text{OH}$  and  $\text{H}_2\text{O}$  for a fixed  $\delta^{18}\text{O}$ – $\text{H}_2\text{O}$ . Overall, while it is difficult to predict the exact range of  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  compositions of  $\text{HNO}_3(1)$  due to its numerous dependences, the proposed mechanism predicts that  $\delta^{15}\text{N}$ – $\text{HNO}_3(1)$  should reflect  $\delta^{15}\text{N}$ – $\text{NO}_2$ , which should be slightly higher than  $\delta^{15}\text{N}$ – $\text{NO}_x$ ; however,

Table 4  
Shift in  $\delta^{15}\text{N}$  and  $\delta^{18}\text{O}$  of  $\text{NO}$ ,  $\text{NO}_2$ , and  $\text{HNO}_3(1)$  relative to  $\text{NO}_x$  as a function of  $f_{\text{NO}_2}$ .

$f_{\text{NO}_2}$	$\delta^{15}\text{N}$ – $\text{NO}$	$\delta^{18}\text{O}$ – $\text{NO}$	$\delta^{15}\text{N}$ – $\text{NO}_2$	$\delta^{18}\text{O}$ – $\text{NO}_2$	$\delta^{15}\text{N}$ – $\text{HNO}_3(1)$	$\delta^{18}\text{O}$ – $\text{HNO}_3(1)$
1	–	–	0	0	0	–57.5
0.85	–30.5	7.0	5.4	–1.2	5.4	–58.3
0.70	–25.3	5.8	10.8	–2.5	10.8	–59.2
0.55	–19.9	4.7	16.3	–3.8	16.3	–60.1



$\delta^{18}\text{O}-\text{HNO}_3(1)$  should be lower than  $\delta^{18}\text{O}-\text{NO}_2$  as a result of the O isotopic mass-balance between  $\text{NO}_2$  and  $\cdot\text{OH}$ .

#### 4.2. Predicted nighttime $\text{HNO}_3$ $\delta^{18}\text{O}-\delta^{15}\text{N}$ space

During the nighttime, higher N oxides form and new pathways of  $\text{HNO}_3$  production become important due to the absence of photochemically produced  $\cdot\text{OH}$ . Under these conditions,  $\text{NO}_2$  is oxidized by  $\text{O}_3$  forming the  $\text{NO}_3$  radical **R5**, which exists at thermal equilibrium with  $\text{NO}_2$  and  $\text{N}_2\text{O}_5$  (Russell et al., 1985; **R6**). Subsequent hydrolysis of  $\text{N}_2\text{O}_5$  on a wetted surface forms  $\text{HNO}_3$ .



This  $\text{HNO}_3$  formation pathway is typically most prevalent during the winter when  $\text{N}_2\text{O}_5$  is thermally stable (Calvert et al., 1985; Chang et al., 2011). Nighttime  $\text{HNO}_3$  may also form through hydrogen abstraction from organic compounds (R) by  $\text{NO}_3$ :



Since photochemical cycling of  $\text{NO}_x$  shuts down during the night, nearly all  $\text{NO}_x$  exists as  $\text{NO}_2$ . Thus,  $\delta^x-\text{NO}_x \approx \delta^x-\text{NO}_2$  as isotopic exchange between  $\text{NO}$  and  $\text{NO}_2$  ceases (Freyer et al., 1993; Walters et al., 2016). However, isotopic equilibrium between  $\text{NO}_2$ ,  $\text{NO}_3$ , and  $\text{N}_2\text{O}_5$  is likely to be achieved mirroring its rapid chemical equilibrium rate (lifetime of 1–2 min; Amell and Daniels, 1952; Freyer, 1991), which will have an impact on  $\delta^{18}\text{O}$  and  $\delta^{15}\text{N}$  of atmospheric nitrate produced during the night. Generally, the tropospheric concentrations of  $[\text{NO}_2] \gg [\text{N}_2\text{O}_5] \gg [\text{NO}_3]$ , as  $\text{NO}_2$  is typically on the order of ppb while  $\text{N}_2\text{O}_5$  and  $\text{NO}_3$  are on the order of ppt (e.g. Chang et al., 2011). Therefore, in most cases the isotopic composition of  $\text{N}_2\text{O}_5$  and  $\text{NO}_3$  should reflect the exchange  $\alpha$ 's with respect to  $\text{NO}_2$  (i.e.  $\alpha_{\text{N}_2\text{O}_5/\text{NO}_2}$  and  $\alpha_{\text{NO}_3/\text{NO}_2}$ ) and can be determined from the following:

$$\delta^x - \text{N}_2\text{O}_5(\text{‰}) = \delta^x - \text{NO}_2(\text{‰}) + 1000[\alpha_{\text{N}_2\text{O}_5/\text{NO}_2} - 1] \quad (15)$$

$$\delta^x - \text{NO}_3(\text{‰}) = \delta^x - \text{NO}_2(\text{‰}) + 1000[\alpha_{\text{NO}_3/\text{NO}_2} - 1] \quad (16)$$

Assuming a negligible kinetic isotopic fractionation associated with **R7**, the isotopic composition of atmospheric nitrate produced through  $\text{N}_2\text{O}_5$  hydrolysis, which we will refer to as  $\text{HNO}_3(2)$ , can be estimated from  $\delta^x-\text{N}_2\text{O}_5$  and O isotopic mass-balance, our calculated exchange fractionation factors, and using an estimated  $\delta^{18}\text{O}-\text{H}_2\text{O}$  range from  $-25$ – $0\text{‰}$ .

$$\delta^{15}\text{N} - \text{HNO}_3(2)(\text{‰}) = \delta^{15}\text{N} - \text{N}_2\text{O}_5(\text{‰}) \quad (17)$$

$$\delta^{18}\text{O} - \text{HNO}_3(2)(\text{‰}) = \frac{5}{6}(\delta^{18}\text{O} - \text{N}_2\text{O}_5(\text{‰})) + \frac{1}{6}(\delta^{18}\text{O} - \text{H}_2\text{O}(\text{‰})) \quad (18)$$

Assuming this model,  $\text{N}_2\text{O}_5$  and  $\text{HNO}_3(2)$  have an estimated  $\delta^{15}\text{N}$  range of  $10.5$ – $25.5\text{‰}$  (Fig. 4b), which is approximately  $25.5\text{‰}$  higher than  $\text{NO}_x$  as a result of the

$\text{N}_2\text{O}_5/\text{NO}_2$  exchange ( $^{15}\alpha_{\text{N}_2\text{O}_5/\text{NO}_2} = 1.0255$  at 300 K).  $\delta^{18}\text{O}-\text{N}_2\text{O}_5$  is estimated to range from  $121.4$ – $131.4\text{‰}$  (Fig. 4b), as a result of the O isotopic exchange between  $\text{N}_2\text{O}_5/\text{NO}_2$  ( $^{18}\alpha_{\text{N}_2\text{O}_5/\text{NO}_2} = 1.0094$  at 300 K). Assuming O mass-balance between  $\text{N}_2\text{O}_5$  and  $\text{H}_2\text{O}$  (Eq. (18)),  $\delta^{18}\text{O}-\text{HNO}_3(2)$  is predicted to range from  $97.0$ – $109.5\text{‰}$  at 300 K. As temperatures decrease, the N and O fractionation from the  $\text{N}_2\text{O}_5/\text{NO}_2$  isotopic exchange will increase, and this should result in higher  $\delta^x-\text{N}_2\text{O}_5$  relative to the local  $\text{NO}_x$ .

Nitrate produced at night formed through the  $\text{NO}_3$  hydrogen abstraction pathway, which we will refer to as  $\text{HNO}_3(3)$ , will have an O and N isotopic composition that is equal to the isotopic composition of  $\text{NO}_3$ , assuming isotopic mass-balance and neglecting any kinetic isotopic fractionation associated with **R8**:

$$\delta^{15}\text{N} - \text{HNO}_3(3)(\text{‰}) = \delta^{15}\text{N} - \text{NO}_3(\text{‰}) \quad (19)$$

$$\delta^{18}\text{O} - \text{HNO}_3(3)(\text{‰}) = \delta^{18}\text{O} - \text{NO}_3(\text{‰}) \quad (20)$$

Assuming this isotopic mechanism,  $\delta^{15}\text{N}$  of  $\text{NO}_3$  and  $\text{HNO}_3(3)$  are estimated to range from  $-33.1$ – $18.1\text{‰}$ , which is approximately  $18.4\text{‰}$  lower than the  $\delta^{15}\text{N}-\text{NO}_2$  (Fig. 4b). Additionally,  $\delta^{18}\text{O}$  of  $\text{NO}_3$  and  $\text{HNO}_3(3)$  is estimated to range from  $102.9$ – $112.9\text{‰}$ , which is approximately  $9.7\text{‰}$  lower than  $\delta^{18}\text{O}-\text{NO}_2$  (Fig. 4b). Both  $\delta^{15}\text{N}$  and  $\delta^{18}\text{O}$  are lower in  $\text{HNO}_3(3)$  relative to  $\text{NO}_2$  due to isotopic exchange between  $\text{NO}_3$  and  $\text{NO}_2$  that favors both  $^{15}\text{N}$  and  $^{18}\text{O}$  partitioning in  $\text{NO}_2$  ( $^{15}\alpha_{\text{NO}_3/\text{NO}_2} = 0.9819$  and  $^{18}\alpha_{\text{NO}_3/\text{NO}_2} = 0.9901$  at 300 K). Our calculations indicate that as temperatures decrease, both the N and O fractionation resulting from the  $\text{NO}_3/\text{NO}_2$  isotopic exchange will increase, which will result in higher  $\delta^x-\text{HNO}_3(3)$  values.

It is important to note the uncertainty in the calculated  $^{15}\beta$  of  $\text{NO}_3$  that is related to the computational difficulties of  $\text{NO}_3$  as previously discussed (e.g. Morris et al., 1990; Dutta et al., 2013). B3LYP/cc-pVTZ, EDF2/cc-pVTZ, and QCISD/cc-pVDZ models yielded  $^{18}\beta$  within  $3.8\text{‰}$  at 300 K (Fig. 2) indicating general agreement. However, these methods found a wider range in  $^{15}\beta$ , from  $1.0715$  to  $1.0864$  at 300 K (Fig. S1), indicating a disagreement as high as  $14.9\text{‰}$ . Despite the uncertainty in  $^{15}\beta$  of  $\text{NO}_3$ , all of the considered computational methods indicate that the  $^{15}\beta$  values of  $\text{NO}_3$  are less than those of  $\text{NO}_2$  (i.e.  $^{15}\alpha_{\text{NO}_3/\text{NO}_2} < 1$ ; Fig. S2). Therefore, while the actual  $^{15}\beta$  values of  $\text{NO}_3$  are somewhat uncertain, our calculations indicate that  $\text{NO}_3$  should have lower  $\delta^{15}\text{N}$  relative to  $\text{NO}_2$ . In our calculations, we have used  $\text{NO}_3$   $^{15}\beta$  and  $^{18}\beta$  values calculated using EDF2/cc-pVTZ, because vibrational frequencies calculated with this method are in closer agreement with experimental data (Kim et al., 1992; Jacox and Thompson, 2008; Beckers et al., 2009) (Table 1).

#### 4.3. Implications of $\delta^{18}\text{O}-\delta^{15}\text{N}$ compositions in $\text{HNO}_3$

Assuming the same range of starting N and O isotopic compositions of  $\text{NO}_x$  ( $\delta^{15}\text{N}-\text{NO}_x$ :  $-15$ – $0\text{‰}$ ,  $\delta^{18}\text{O}-\text{NO}_x$ :  $112$ – $122\text{‰}$ ), our calculations suggest that different  $\text{HNO}_3$  formation pathways yield relatively distinctive  $\delta^{18}\text{O}-\delta^{15}\text{N}$  compositions (Fig. 5). Our calculations indicate that

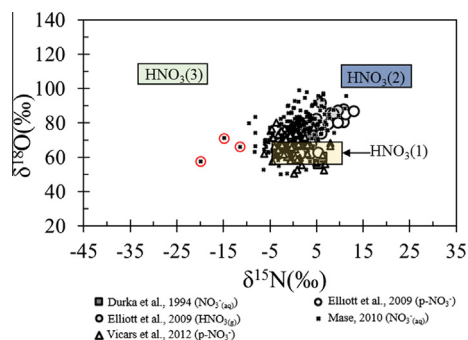


Fig. 5. Predicted  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  compositions for three major  $\text{HNO}_3$  production pathways that is compared with previous atmospheric nitrate measurements (Durka et al., 1994; Elliott et al., 2009; Mase, 2010; Vicars et al., 2013). Where  $\text{HNO}_3(1)$ ,  $\text{HNO}_3(2)$ , and  $\text{HNO}_3(3)$  represent various  $\text{HNO}_3$  production pathways that include  $\text{NO}_2 + \cdot\text{OH} \rightarrow \text{HNO}_3(1)$ ,  $\text{N}_2\text{O}_5 + \text{H}_2\text{O} + \text{surface} \rightarrow 2\text{HNO}_3(2)$ , and  $\text{NO}_3 + \text{R} \rightarrow \text{HNO}_3(3) + \text{R}'$ . The data points within the red circles are outside of the predicted equilibrium  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  space, which may indicate a  $\text{NO}_x$  source with an extremely low  $\delta^{15}\text{N}$  value such as soil denitrification. (For interpretation of the references to colour in this figure legend, the reader is referred to the web version of this article.)

$\text{HNO}_3(1)$  should result in the lowest  $\delta^{18}\text{O}$  values and mid-ranged  $\delta^{15}\text{N}$  values. Our model predicts that  $\text{HNO}_3(2)$  and  $\text{HNO}_3(3)$  should have high  $\delta^{18}\text{O}$  values that are similar. However,  $\text{HNO}_3(2)$  has the highest predicted  $\delta^{15}\text{N}$  values, while  $\text{HNO}_3(3)$  has the lowest  $\delta^{15}\text{N}$  values. This suggests that  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  arrays for different  $\text{HNO}_3$  production pathways can be used as isotopic end-members in isotope mixing models that might explain  $\delta^{15}\text{N}$  and  $\delta^{18}\text{O}$  values observed in atmospheric nitrate.

These predicted  $\delta^{15}\text{N}$  and  $\delta^{18}\text{O}$  isotopic end-members were compared with measured  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  values of atmospheric nitrates (Fig. 5), including  $\text{HNO}_3(\text{g})$  (Elliott et al., 2009),  $\text{p-NO}_3^-$  (Elliott et al., 2009; Mase, 2010; Vicars et al., 2013), and  $\text{NO}_3^-(\text{aq})$  (Durka et al., 1994). The measured  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  values in atmospheric nitrate plot between the two major  $\text{HNO}_3$  formation pathways  $\text{HNO}_3(1)$  and  $\text{HNO}_3(2)$  predicted by this equilibrium model (Fig. 5). However, some of the reported  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  values appear to represent a mixture of  $\text{HNO}_3(1)$  and  $\text{HNO}_3(3)$ , which is most likely to occur for areas with high DMS/VOC concentrations such as marine and forest ecosystems (Geyer and Platt, 2002). Our predicted  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  compositions may explain both the  $\delta^{18}\text{O}$  and  $\delta^{15}\text{N}$  seasonal cycling typically observed in atmospheric nitrate, in which  $\delta^{18}\text{O}$  and  $\delta^{15}\text{N}$  values are highest in the winter and lowest during the summer (e.g. Freyer, 1978, 1991; Elliott et al., 2009; Mase, 2010; Beyn et al., 2014). During winter, the  $\text{HNO}_3(2)$  pathway is generally most prominent (Calvert et al., 1985; Chang et al., 2011), and our calculations suggest that this pathway results in high  $\delta^{18}\text{O}$  and  $\delta^{15}\text{N}$  values. During summer, the  $\text{HNO}_3(1)$  pathway is generally most prominent due to the higher concentrations of photochemically produced  $\cdot\text{OH}$  (Calvert et al., 1985), and our calculations suggest this pathway will result in  $\delta^{15}\text{N}$

and  $\delta^{18}\text{O}$  values lower than those for the  $\text{HNO}_3(2)$  pathway.

While  $\delta^{18}\text{O}$  and  $\Delta^{17}\text{O}$  of atmospheric nitrate has been previously linked to  $\text{NO}_x$  oxidation pathways (e.g. Michalski et al., 2003, 2012; Savarino et al., 2007), our calculations suggest that  $\delta^{15}\text{N}$  may also be linked to oxidation chemistry. However, seasonal changes in  $\delta^{15}\text{N}$  of atmospheric nitrate will also be impacted by changes in seasonal emissions of  $\text{NO}_x$  by a variety of processes, particularly  $\text{NO}$  emitted by nitrification and denitrification occurring in soils. In areas where  $\text{NO}_x$  emissions should not change seasonally, such as areas dominated by fossil-fuel combustion, any observed seasonal change in  $\delta^{15}\text{N}$  of atmospheric nitrate suggests seasonal changes in  $\text{NO}_x$  oxidation chemistry. In areas in which  $\text{NO}_x$  emissions exhibit a large seasonal change, such as agricultural regions, the  $\delta^{15}\text{N}$  of atmospheric nitrate will be a function of both the seasonal change in  $\delta^{15}\text{N}$ – $\text{NO}_x$  and in the  $\text{NO}_x$  oxidation equilibrium isotope effect. The equilibrium model indicates that it may be difficult to partition  $\text{NO}_x$  sources solely from  $\delta^{15}\text{N}$ – $\text{HNO}_3$  values, since the  $\delta^{15}\text{N}$  tends to be sensitive to equilibrium effects that may alter the original  $\text{NO}_x$  source  $\delta^{15}\text{N}$  value as it is oxidized to  $\text{HNO}_3$  (Fig. 4). However, evaluation of  $\delta^{15}\text{N}$ – $\delta^{18}\text{O}$  space of atmospheric nitrate may help elucidate  $\text{NO}_x$  emission sources in some cases. Atmospheric nitrate outside of the calculated  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  compositional range might indicate its  $\text{NO}_x$  precursor had a distinctive  $\delta^{15}\text{N}$  outside of the typical range of  $-15$  to  $0\text{‰}$ . For example, soil denitrification events, which are estimated to have a low  $\delta^{15}\text{N}$ – $\text{NO}_x$  (approximately between  $-50$  and  $-20\text{‰}$ ; Li and Wang, 2008; Felix and Elliott, 2013, 2014) may be trackable with atmospheric nitrate  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  measurements. If emitted  $\text{NO}_x$  is oxidized through the  $\text{HNO}_3(1)$  pathway, it should have low  $\delta^{15}\text{N}$  and  $\delta^{18}\text{O}$ , which does not match the equilibrium predictions (Fig. 5). This sort of event may explain why a few of the atmospheric nitrate measurements from Mase (2010) is outside of the general predicted  $\text{HNO}_3$  formation  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  range (Fig. 5). These rainwater nitrates were collected at N deposition sites after a storm saturated recently fertilized agricultural fields in the Midwestern U.S., suggesting  $\text{NO}$  produced by nitrification/denitrification.

While our predicted equilibrium  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  compositions generally agrees with atmospheric nitrate measurements (e.g. Durka et al., 1994; Elliott et al., 2009; Mase, 2010; Vicars et al., 2013), it is important to point out the simplicity of our model. Primarily, we have ignored any kinetic isotope effects. If equilibrium between  $\text{NO}_2$ ,  $\text{NO}_3$ , and  $\text{N}_2\text{O}_5$  is achieved, kinetic isotope effects associated with the formation of  $\text{NO}_3$  and  $\text{N}_2\text{O}_5$  should be erased. However, kinetic isotope effects will need to be considered for the photochemical cycling of  $\text{NO}_x$ , that includes  $\text{NO}_2$  photolysis and  $\text{NO}$  oxidation, as these MDFPs are predicted to have an impact on the N and O isotopic composition of  $\text{NO}_2$  (Walters et al., 2016), which may then be propagated into  $\text{NO}_y$  molecules. Additionally, we will need to determine kinetic isotope fractionation factors for the final step in forming atmospheric nitrate (i.e. R4, R7, and R8), which may play an important role in its N and O isotopic compo-

sition. Our equilibrium model suggests that the isotopic composition of  $\text{NO}_2$  drives the isotopic composition of atmospheric nitrate indicating the need for future *in situ* isotopic measurements of  $\text{NO}_2$ . Determination of  $\text{NO}_x$  photochemical cycling fractionation factors, isotopic *in situ* measurements of  $\text{NO}_2$ , and kinetic isotopic modeling of  $\text{NO}_x$  oxidation will be the subject for future research.

## 5. CONCLUSIONS

Hybrid density functional theory calculations indicate that oxygen equilibrium isotopic exchange involving  $\text{NO}_y$  molecules generally favors the formation  $^{18}\text{O}$  in  $\text{NO}_y$  molecules other than  $\text{NO}$  and  $\text{NO}_2$  at the most likely O exchange sites. Therefore, equilibrium isotope exchange may play a role in the  $\delta^{18}\text{O}$  of atmospheric nitrate in addition to the O isotopic mass-balance of its precursor molecules and oxidants. A simple equilibrium and mass-balance model indicates that the three major  $\text{HNO}_3$  formation pathways may have unique  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  compositions. This model generally predicts a range of  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  values consistent with  $\text{HNO}_3$  measurements and suggests that  $\delta^{18}\text{O}$ – $\delta^{15}\text{N}$  space of atmospheric nitrate may provide useful information about the conditions of  $\text{NO}_x$  oxidation pathways. This model may explain the seasonal changes typically observed in both  $\delta^{18}\text{O}$  and  $\delta^{15}\text{N}$  of atmospheric nitrate.

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## APPENDIX A. SUPPLEMENTARY DATA

Supplementary data associated with this article can be found, in the online version, at <http://dx.doi.org/10.1016/j.gca.2016.06.039>.

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